Radio-optical outliers – a case study with ICRF2 and SDSS

Sándor Frey FÖMI Satellite Geodetic Observatory & MTA Research Group for Physical Geodesy and Geodynamics (Budapest, Hungary)

Gábor Orosz Department of Geodesy and Surveying, Budapest University of Technology and Economics







QSO Astrophysics, Fundamental physics, and Astrometric Cosmology in the Gaia era GREAT-ESF Workshop, 6–9 June 2011, Porto, Portugal

Outline of the talk

Direct link between the best radio and optical reference frames astrometric and astrophysical importance

Are the common objects really coincident? no Gaia yet – a case study with identifying ICRF2 sources in SDSS

How accurate is the SDSS astrometry in general? Data Releases 7 & 8

Individual objects that are significantly non-coincident *how many and why?*



Radio-optical link

Radio reference frame: based on VLBI positions of compact extragalactic radio sources (active galactic nuclei) ICRF2 - Fey, Gordon & Jacobs (eds.) 2009

The quasars are too faint in optical

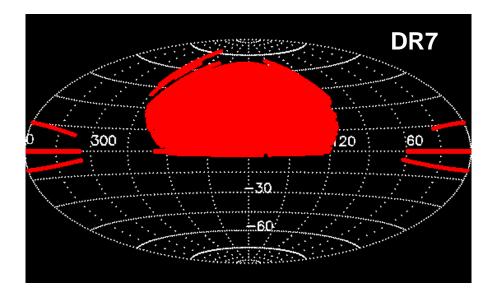
with the current catalogues, the radio-optical link is possible only indirectly, using a couple of radio-emitting stars problems: quite few of them; proper motion inaccuracies deteriorate the quality with time

Gaia: a direct alignment of the reference frames – for the first time ~half a million quasars with a limiting magnitude of $m_V \sim 20$ thousands of "good" bright sources will define an accurate frame common Gaia–VLBI quasars \rightarrow direct link

Are the optical and radio positions identical? not necessarily; the real question is how different they are



While waiting for Gaia...



Sloan Digital Sky Survey (SDSS) Data Release 7 (DR7) cas.sdss.org/dr7

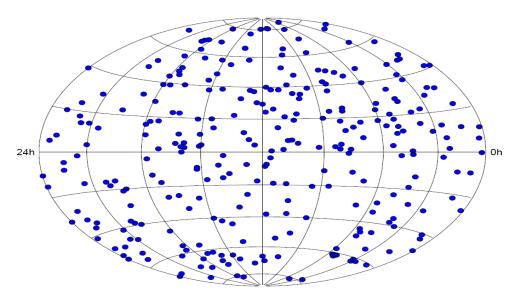
Data Release 8 (DR8) skyserver.sdss3.org/dr8

coverage: over 1/3 of the sky magnitude limit: ~22^m

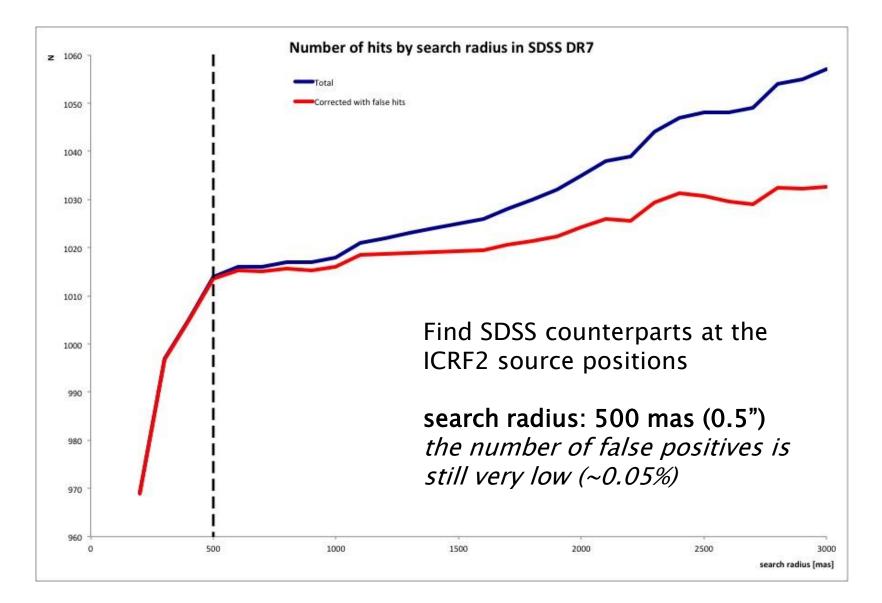
2nd realization of the International Celestial Reference Frame (ICRF2)

IERS Technical Note 35

295 defining sources, 3414 objects in total

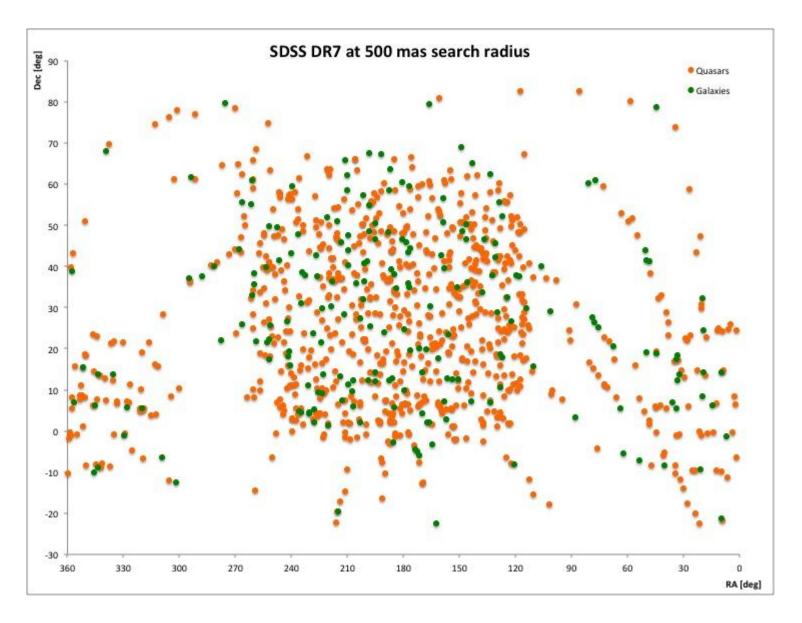


ICRF2–SDSS cross-identifications



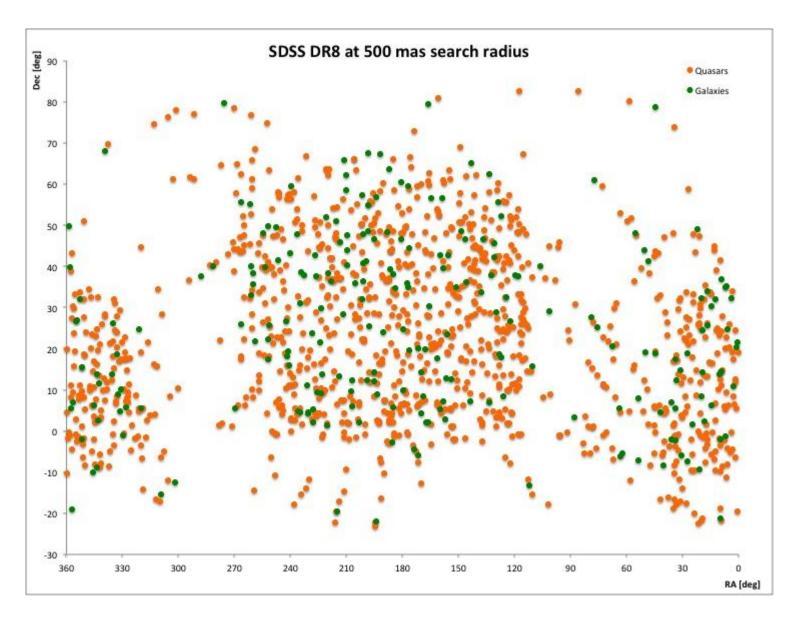


Common objects on the sky (DR7)



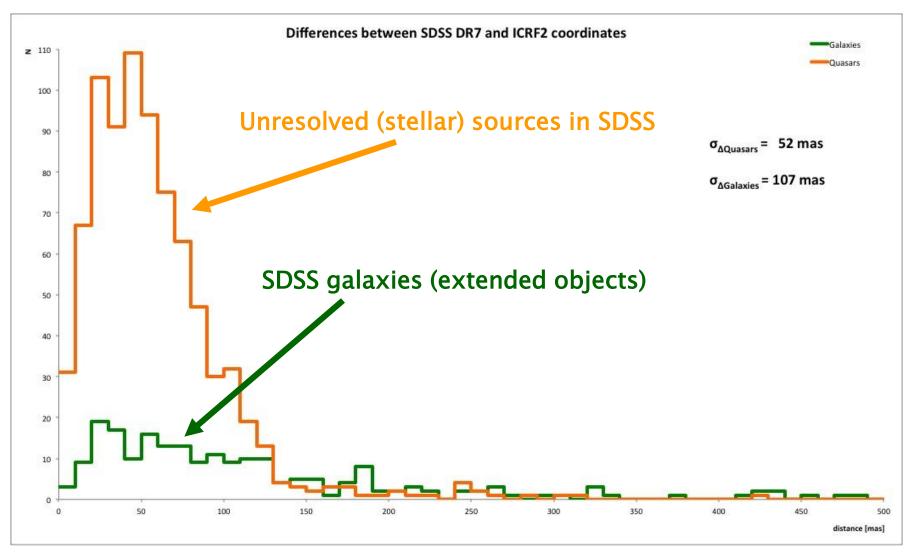


Common objects on the sky (DR8)



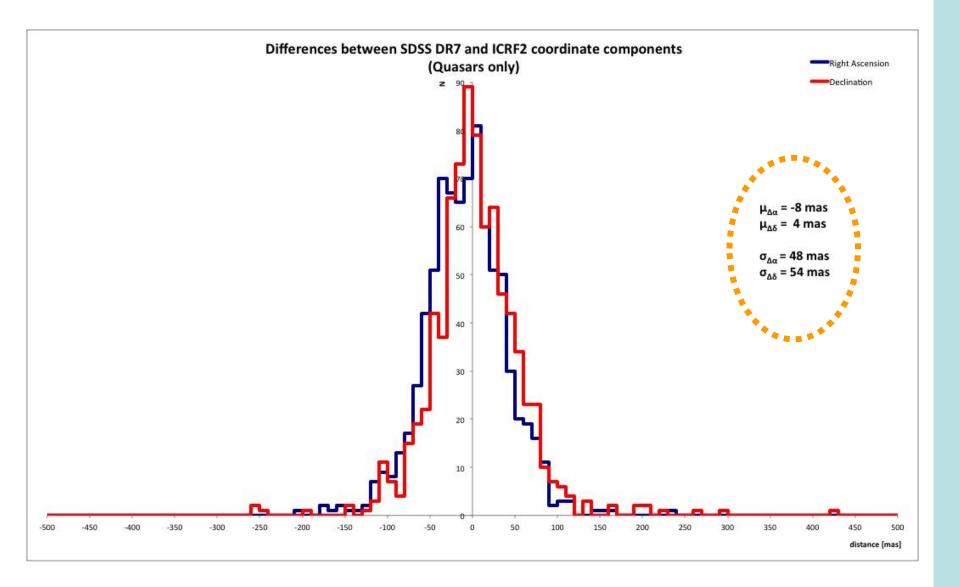


The distribution of SDSS position offsets

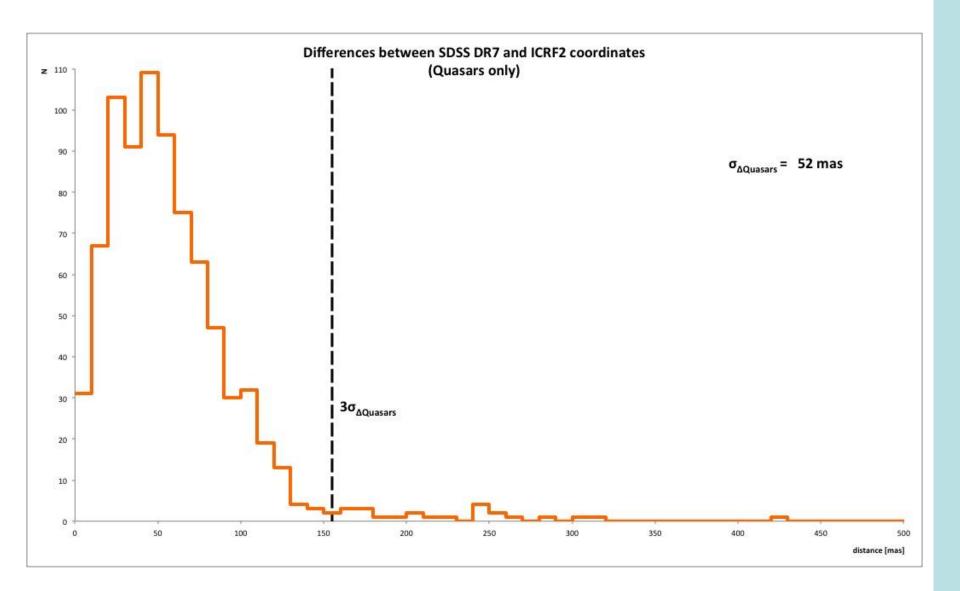


The stated astrometric accuracy of SDSS DR7 is <100 mas in each coordinate

Kömi (

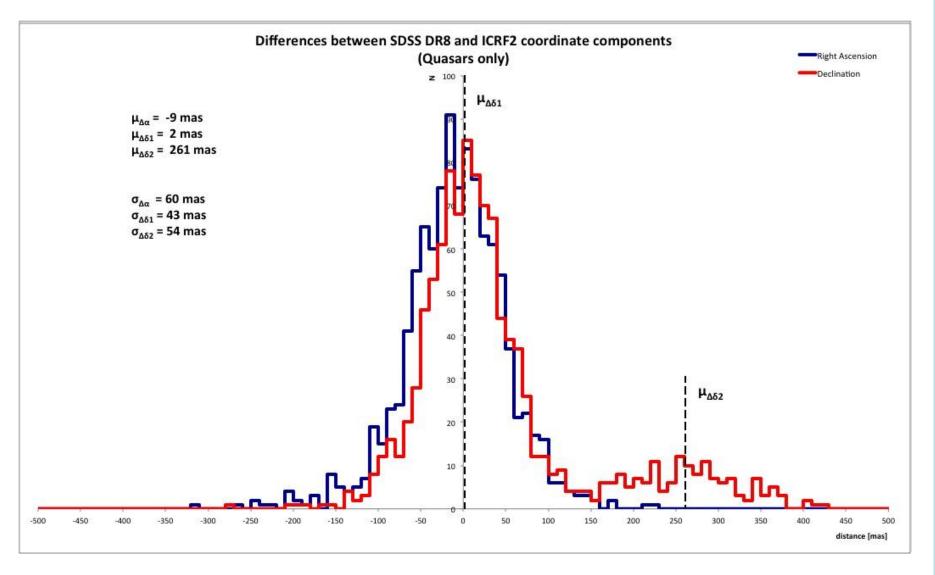


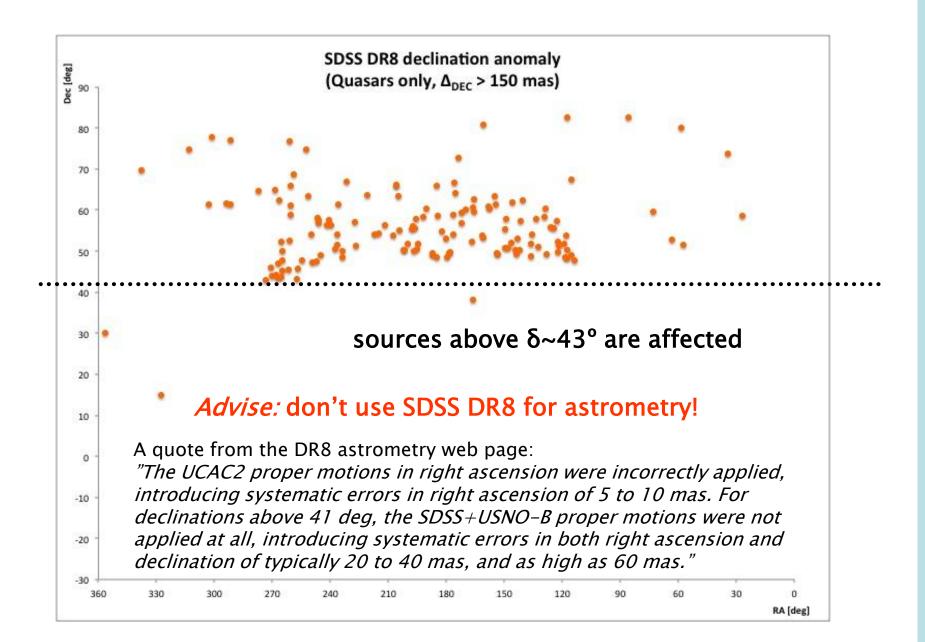
<u>Kömi</u>



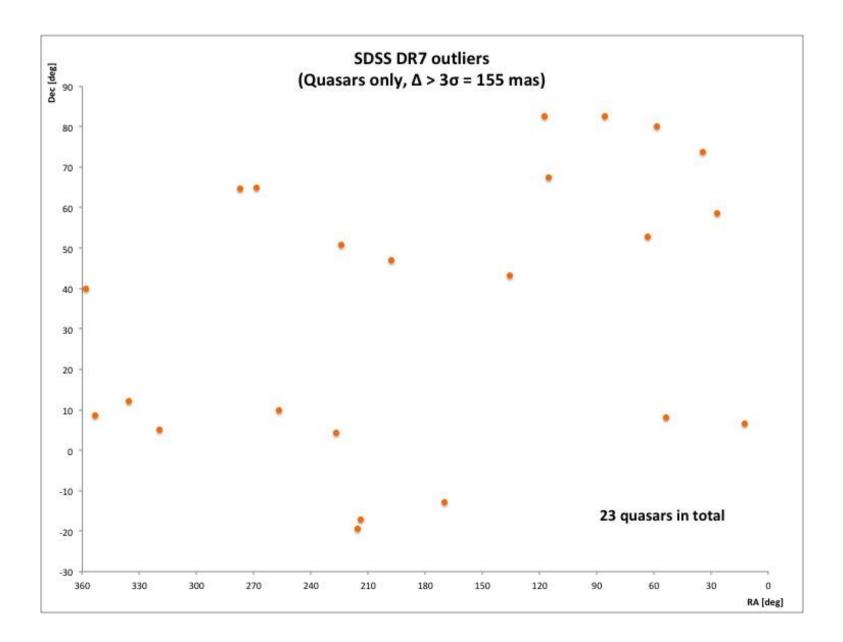


A quick look at DR8...











Outliers in DR7

There are ~20 SDSS optical quasars with at least 3σ (155 mas) positional offset from the corresponding ICRF2 position

Are they chance identifications?

Certainly not! Simulations using a large "fake" ICRF catalogue show that the probability of chance coincidence is 0.02% within 300 mas, and 0.05% within 500 mas \rightarrow only 1 or 2 unrelated optical sources are expected within our search radius

Do they result from astrometric mis-calibration in SDSS?

Maybe... This cannot be excluded, at least for some of the sources / SDSS fields. But probably does not explain all outliers.



Astrophysical reasons?

Core shift?

Opacity effect (frequency-dependent position of the VLBI "core") ~0.1 mas systematic difference between the radio (8 GHz) and optical Andrei Lobanov's and Kirill Sokolovsky's talks e.g. Sokolovsky et al. 2011 (A&A, in press) Kovalev et al. 2008 (A&A 483, 759) --- too small compared to our offsets

Extended vs. compact radio sources?

Optical-radio position differences in a large sample of radio sources are systematically higher (~8 mas) for extended radio sources da Silva Neto et al. 2002 (AJ 124, 612)

--- too small compared to our offsets



Astrophysical reasons? (cont.)

Gravitational lensing?

The compact radio emission originates from a background source, while the optical is dominated by the lensing foreground galaxy lan Browne's and Francois Finet's talks --- how many of them? probably can't explain each outlier

Dual AGNs?

The (projected) linear size corresponding to 150–300 mas angular size at the redshift of e.g. z=1 is ~1.2–2.4 kpc (using the "standard" cosmological model) These correspond to sub-galactic sizes ––– mergers with such separations do exist

Radio emission from e.g. double-peaked narrow [O III] emission line quasar pairs from SDSS is detected in ~50 cases, although these are generally weak radio sources – unlike ICRF2 objects (Li C. et al. 2011, in prep.)



Summary

> There are 1014 ICRF2 radio sources with optical counterparts in SDSS DR7, the majority of them are quasars (i.e. optically unresolved)

SDSS DR7 quasar positions are accurate to ~50 mas

> Don't touch SDSS DR8 if you have astrometry in mind...

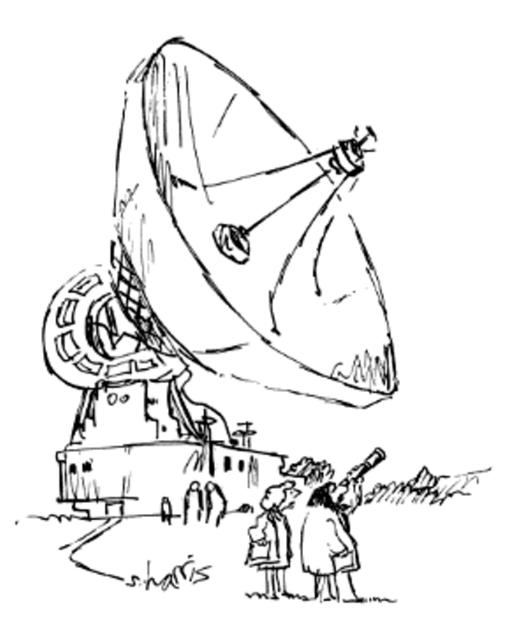
> There are over 20 quasars for which the ICRF2 and SDSS positions exceed 155 mas (3 σ)

> These should be further investigated to reveal the cause(s) of such large non-coincidences

When the accurate optical and radio frames are finally aligned, one must filter out such peculiar sources from the link objects (and study them from astrophysical points of view!)



Thank you for your attention!



"Just checking."



Radio-optical outliers...