Putting Einstein to Test Astrometric Experiments in Fundamental Physics

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- Modelling the observations

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- The quadrupole component of the gravitational field
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 $\begin{array}{c} \textbf{Theoretical background}\\ \textbf{Estimation of the PPN } \gamma \text{ parameter}\\ \textbf{Estimation of the PPN } \beta \text{ parameter}\\ \textbf{Beyond the PPN framework} \end{array}$

Astrometry and gravity Modelling the observations

Why Astrometry

First experimental evidences of General Relativity came from Astrometry.

• One can thus hope that the discussion of the observations of Mercury simply confirm previous research. Now this is not negligible: we see here that the approximately 3-fold secular movement of the eccentricity, added to the secular movement of the perihelion, gives a sum in which the observations are greater by 39" than those which result from calculation (Le Verrier, 1859)

 Thus the results of the expeditions to Sobral and Principe can leave little doubt that a deflection of light takes place in the neighbourhood of the Sun and that it is of the amount demanded by Einstein's generalised theory of relativity, as attributable to the sun's gravitational field. (Dyson, Eddington and Davidson, 1920)

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Astrometry and gravity Modelling the observations

The PPN Framework

The Parametrized Post-Newtonian formalism is a framework which has been designed to describe the Post-Newtonian limit of all the metric theories of gravity by means of a set of 10 parameters. These parameters are associated to specific physical properties.

	How much space-curvature	(/ a	(12	
	produced by unit rest mass?	~3		
	How much "nonlinearity" in the		Violation of conservation of total	
	superposition law for gravity?	51	momentum?	
	Prefered-location effects?			
α_1	Prefered-frame effects?	ζз		
		ζ4		
α3				

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Parameter	Meaning	Parameter	Meaning
γ	How much space-curvature	<i></i>	
	produced by unit rest mass?	u3	
β	How much "nonlinearity" in the	×	Violation of conservation of total
	superposition aw for gravity?	51	momentum?
ξ	Prefered-location effects?	ζ2	
α1		ζ3	
α2	Prefered-frame effects?	ζ4	
α3			

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Astrometry and gravity Modelling the observations

Theoretical Implications

A precise estimation of the PPN γ and β parameters, and in general precise tests of fundamental physics, have theoretical and observational implications.

ulfilling theoretical needs

They are the phenomenological "trace" of a scalar field coupled with gravity which is related to:

- theories aiming to provide a formulation of a quantum theory of gravity;
- theories fully compatible with the Mach principle;
- cosmological scenarios with inflationary stage.

planning, observational evidences.

They can provide constraints on theories which claim of being able to account for several astrophysical and cosmological problems without any need for DM or DE like. e.g.:

- acceleration of cosmological expansion (attributed to DE);
- galactic rotation curves, galaxy cluster masses and dynamics (attributed to DM);
- observational data from gravitational lensing;
- Tully-Fisher relation.

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Astrometry and gravity Modelling the observations

Present Experimental Limits

Present	best	resu	ts	for	γ	
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Experiment	Effect	Technique	$ \gamma - 1 $ lower bound	
HIPPARCOS	Light	Global	2 10-3	
	deflection	Astrometry	5.10	
VIBI	Light	Radio	4 5 10-4	
VLDI	deflection	Interferometry	4.5 * 10	
	Shapiro	Round-trip		
Cassini	time delay	travel time of	$2.3 \cdot 10^{-5}$	
	time delay	radar signals		

References: Froeschlé et al. (1997), Shapiro et al. (2004), Bertotti et al. (2003).

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Astrometry and gravity Modelling the observations

Present Experimental Limits

Experimental bounds on β depends on the estimation of γ .

Present best results for eta				
Data	Effect (Technique	Combination	meta-1 lower	
Data	Effect/Technique	with γ	bound	
Radar obs. of	Shift excess for	Day B	2 10-3	
Mercury (1966-1990)	Mercury	$2\gamma - p$	3.10	
Lunar Laser Ranging	Nordtvedt effect	$4eta-\gamma$	$\sim 10^{-4}$	
Radar obs. of inner	"Grand fits" of Solar			
bodies	System eqs. of	$2\gamma - \beta$	$1 \div 2 \cdot 10^{-4}$	
(1963-2003)	motion			

References: Shapiro (1990), Pitjeva (2005, 2010), Williams et al. (2004), Will (2006).

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Astrometry and gravity Modelling the observations

The Observable in the Euclidean World

- The basic astrometric observable is an angle between two viewing directions $\cos \psi = \frac{\mathbf{r}_1 \cdot \mathbf{r}_2}{|\mathbf{r}_1| |\mathbf{r}_2|}$ (1)
- We need to find a relativistic equivalent of this formula.



Astrometry and gravity Modelling the observations

Theory of Measurements for the Impatient

- A physical observer is identified by his/her worldline Γ, whose four-velocity is u^α ≡ Γ (⇒the parameter s of Γ represents the proper time of the observer).
- \forall point $\in \Gamma$, u^{α} splits the space-time into a 1D subspace $|| u^{\alpha}$ and a 3D subspace $\perp u^{\alpha}$, which are the *time direction* and the *space* relative to the observer u^{α} respectively.

$$\mathscr{P}_{\alpha\beta} = -u_{\alpha}u_{\beta}$$

$$_{\alpha} u_{\beta}$$
 Parallel projector

$$\mathscr{T}_{\alpha\beta} = g_{\alpha\beta} + u_{\alpha}u_{\beta}$$



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$$ds^{2} = g_{\alpha\beta}dx^{\alpha}dx^{\beta}$$

= $-c^{2}\underbrace{\left(-c^{-2}\mathscr{P}_{\alpha\beta}dx^{\alpha}dx^{\beta}\right)}_{dt_{u}^{2}} + \underbrace{\left(\mathscr{T}_{\alpha\beta}dx^{\alpha}dx^{\beta}\right)}_{dt_{u}^{2}}$

A. Vecchiato Astrometric Experiments in Fundamental Physics

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Astrometry and gravity Modelling the observations

A Relativistic formula for the Astrometric Observable



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Gaia and the global sphere reduction A different approach: relative astrometry

The Concept of the Global Sphere Reconstruction The ideal picture





$$N_{\rm unk} = 4$$

$$N_{\rm arcs} = 1$$

A. Vecchiato Astrometric Experiments in Fundamental Physics

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$$N_* = 3$$
$$N_{\rm unk} = 6$$

$$N_{\rm arcs} = 3$$

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The Concept of the Global Sphere Reconstruction The ideal picture





$$N_{\rm arcs} = 6$$

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 $N_{\rm arcs} = 10$

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Gaia and the global sphere reduction A different approach: relative astrometry

The Concept of the Global Sphere Reconstruction The (almost) real picture

- To take into account of observational errors:
 - overdetermined system of equations;
 - solution in the least-squares sense.



Gaia and the global sphere reduction A different approach: relative astrometry

The Relativistic Observable for Gaia

The relativistic observable takes into account that Gaia measures abscissae along the scanning direction. E_{3}

$$\cos \Psi_{(\hat{a},k)} = \frac{\mathscr{T}_{\mu\nu} E^{\mu}_{\hat{a}} k^{\nu}}{\sqrt{\mathscr{T}_{\mu\nu} k^{\mu} k^{\nu}}}$$

$$\cos\xi = \frac{\cos\psi_{(\hat{1},k)}}{\sqrt{1 - \cos^2\psi_{(\hat{3},k)}}}$$



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Gaia and the global sphere reduction A different approach: relative astrometry

Estimated Accuracy on γ for Gaia

• In principle, each observation is a function of Astrometric, Attitude, Instrument, and Global parameters.

$$k^{v} = k^{v} (\alpha_{*}, \delta_{*}, \pi_{*}, \mu_{\alpha *}, \mu_{\delta *}, \gamma)$$

$$E_{\hat{a}}^{v} = E_{\hat{a}}^{v} \left(q_{1}^{(a)}, q_{2}^{(a)}, q_{3}^{(a)}, q_{3}^{(a)}, \gamma\right)$$

$$\cos \psi_{(\hat{a},k)} \equiv F_{\hat{a}} \left(\underbrace{\alpha_{*}, \delta_{*}, \pi_{*}, \mu_{\alpha *}, \mu_{\delta *}}_{\text{Astrometric parameters}}, \underbrace{q_{1}^{(a)}, q_{2}^{(a)}, q_{3}^{(a)}, q_{4}^{(a)}}_{\text{Attitude parameters}}, \underbrace{\gamma, \dots}_{\text{Global}}\right)$$

- The dependence on γ gives the estimation of this parameter as a by-product of the sphere reconstruction.
- Current simulations suggest a final accuracy of $\sigma_\gamma \sim 10^{-6}.$

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Gaia and the global sphere reduction A different approach: relative astrometry

Problems and Pitfalls

- Dimension and complexity of the problem ($\sim 10^{10} \times 10^8$)
 - Is convergence ensured?
 - Is the right convergence ensured?
- Correlation with parallaxes
- Correlation with Basic Angle variations



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Gaia and the global sphere reduction A different approach: relative astrometry

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Gaia and the global sphere reduction A different approach: relative astrometry

The GAME Concept

GAME: a concept for a dedicated mission of relative astrometry



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Gaia and the global sphere reduction A different approach: relative astrometry

Estimating γ with GAME

Magnitude	σ _{1 00 s} (mas)	σ _{500s} (mas)
10	0.058	0.026
11	0.093	0.041
12	0.150	0.065
13	0.250	0.103
14	0.423	0.164
15	0.782	0.263
16	1.598	0.425
17	3.581	0.704
18	8.523	1.225

$$\Delta \psi = \frac{(1+\gamma) \, GM_{\odot}}{c^2 \, r_{\rm o}} \sqrt{\frac{1+\cos\psi}{1-\cos\psi}} \ \Rightarrow \ \frac{\sigma_{\Delta\psi}}{\Delta\psi} \sim \frac{\sigma_{\gamma}}{\gamma}$$

$$\Delta \psi \sim 0'' 2$$
 at $\psi = 2^\circ \Rightarrow \frac{\sigma_{\gamma}}{\gamma} \sim 10^{-3}$



Light Deflection

Physical phenomena associated with β Estimation of β with Gaia Estimation of β with GAME

• The β parameter enters as a second-order term into the light deflection

$$\Delta \alpha = f\left(\left(g_{00}\right)^{-1}, g_{\mu\nu}\right)$$

$$egin{aligned} g_{00} &= -1 + 2 U_\odot - 2eta \, U_\odot^2 \ g_{ii} &= 1 + 2 \gamma U_\odot \end{aligned}$$

• This is the reason why the light deflection is an intrinsically bad method to estimate the β parameter.

Physical phenomena associated with β Estimation of β with Gaia Estimation of β with GAME

The Nordtvedt Effect

• The Nordtvedt effect refers to the possibility that massive bodies fall with different accelerations depending on their gravitational self-energy, thus violating the weak equivalence principle.

$$\begin{aligned} \mathbf{a} &= \frac{M_{\mathrm{P}}}{M} \nabla U \\ &\frac{M_{\mathrm{P}}}{M} = 1 - \eta \frac{E_{\mathrm{g}}}{M} \\ \eta &= 4\beta - \gamma + f\left(\xi, \alpha_1, \alpha_2, \zeta_1, \zeta_2\right) \end{aligned}$$

• It is measured with LLR techniques, not with astrometry.

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Physical phenomena associated with β Estimation of β with Gaia Estimation of β with GAME

Perihelion Shift Excess

 The effect of the perihelion shift excess depends on a combination of both γ and β (and other PPN parameters).



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$$\Delta \omega = \frac{6\pi m}{a(1-e^2)} \left[\frac{1}{3} (2+2\gamma-\beta) + f(\alpha_1, \alpha_2, \alpha_3, \zeta_2, J_2) \right]$$

Physical phenomena associated with β Estimation of β with Gaia Estimation of β with GAME

Integration of NEOs Orbits

- Gaia will not be able to estimate the value of β from Mercury (too close to the Sun).
- It will instead revert to the observation of Near-Earth Objects, which are tipically in less favourable conditions, but their number is large ($\sim 10^5$).
- Determining the relativistic orbital corrections for several objects at different inclinations gives also the opportunity of estimating also the solar J_2 (but beware of corr $(\beta, J_2) > 0.9!$).
- Numerical simulations have suggested formal accuracies of

$$\sigma_eta\simeq 1.5\cdot 10^{-4} \qquad \sigma_{J_2}\simeq 5\cdot 10^{-9}$$

using the Gaia observations for some 1300 asteroids.

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Physical phenomena associated with β Estimation of β with Gaia Estimation of β with GAME

Estimated Accuracy

- The shift excess for Mercury is $\Delta \omega \simeq 0.104''/orbit$
- At opposition $V_{\rm merc} \gtrsim -1.5$, $\sigma_{\alpha} \sim 0.5 \, {\rm mas}$ (formal) for a single exposure of 0.1s, i.e.

$$\begin{array}{l} \Delta \alpha \left\{ \begin{array}{c} \simeq 0.27 \Delta \omega \simeq 28 \ \mathrm{mas} \\ = k \Gamma \ (\Gamma \equiv 2 + 2\gamma - \beta) \\ \sigma_{\Delta \alpha} \simeq 0.5 \ \mathrm{mas} \end{array} \right\} \Rightarrow \frac{\sigma_{\beta}}{\beta} \simeq 3 \frac{\sigma_{\Gamma}}{\Gamma} = 3 \frac{\sigma_{\Delta \alpha}}{\Delta \alpha} \sim 5 \cdot 10^{-1} \end{array}$$

 Therefore, one day per orbit (about 88 days) with Mercury at opposition means (assuming ~3 orbits per year)

Years	# of obs. 10 ⁻⁷	$\left(\sigma_{\beta}/\beta\right)\cdot 10^{6}$
1	9.5	5.5
2	19.8	3.9
3	28.4	3.2
4	37.8	2.8
5	47.8	2.5



The J₂ component Further tests of fundamental physics

Astrometric Effect of J_2 : the q-effect

A non-spherically symmetric distribution of the mass in a perturbing body, according to GR, induces specific patterns in the nearby light deflection, which can be modeled w.r.t. the higher order multipoles of its gravity field (q-effect). General Relativity ⇒ ε = 1.

$$\Delta \Phi = \Delta \Phi_1 \mathbf{n} + \Delta \Phi_2 \mathbf{m}$$

$$\Delta \Phi_{1} = \frac{2(1+\gamma)M}{b} \left[1 + \varepsilon J_{2} \frac{R^{2}}{b^{2}} \left(1 - 2(\mathbf{n} \cdot \mathbf{z})^{2} - (\mathbf{t} \cdot \mathbf{z})^{2} \right) \right]$$

$$\Delta \Phi_2 = \frac{4(1+\gamma)M\varepsilon J_2R^2\gamma}{b^3} (\mathbf{m} \cdot \mathbf{z})(\mathbf{n} \cdot \mathbf{z})$$

- These effects has never been measured ⇒ new GR tests (first order at which GW can be modelled).
- Such kind of experiments are foreseen for Gaia (Crosta & Mignard, 2006).
- GAME, as a pointed mission, is able to select the most favourable observing conditions for this measure.



The J_2 component Further tests of fundamental physics

Estimating the q-effect

The Differential Approach

- This method exploits the fact the q-effect has a limited duration as Jupiter moves "away" from its closest approach to suitable stars, comparing the configuration of a stellar field with and without a perturbing body in between.
- This approach requires a favorable configuration and for Gaia it depends on the initial conditions of the Scanning Law (up to 10σ-level configurations with the present mission parameters), while GAME in principle is more flexible.

The Global Approach

- Only in the case of Gaia, another possibility is to introduce the quadrupole contribution to the light deflection into the model for the Gaia abscissa.
- In this way ε, similarly to γ, can be obtained as a by-product of the sphere reduction.

$$\mathbf{n} = \mathbf{\sigma} + \delta \sigma_{pN} + \delta \sigma_Q$$

$$\delta\sigma_Q = \varepsilon\sigma \times \Delta_Q \dot{x}_p(t_{\rm obs}) \times \sigma$$

The J_2 component Further tests of fundamental physics

The Mansouri-Sexl Formalism

This is about gravity INDEED!

- **Special Relativity** requires that *c* is the same for any observer.
- Lorentz transformations are the consequence of this hypothesis.
- Local Lorentz Invariance violation is foreseen in some quantum gravity/high energy theories.
- The Mansouri-Sexl formalism, similarily to the PPN one, generalizes the Lorentz transformations, introducing a set of arbitrary parameters (a, b, d, ε) (Laemmerzahl, 2006)

$$T = \frac{(t - \varepsilon \cdot \mathbf{x})}{a}$$
$$X = \frac{\mathbf{x}}{d} - \left(\frac{1}{d} - \frac{1}{b}\right) \frac{\mathbf{w} \cdot \mathbf{x}\mathbf{w}}{\mathbf{w}^{2}} + \mathbf{w}T$$

• From the astrometric point of view, the LLI violation shows itself as an aberration effect that depends on f = d/b, and puts to test the same properties of the Michelson-Morley experiment (Klioner, 2008).

The J_2 component Further tests of fundamental physics

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